



Status of S-NPP VIIRS Solar and Lunar Calibration

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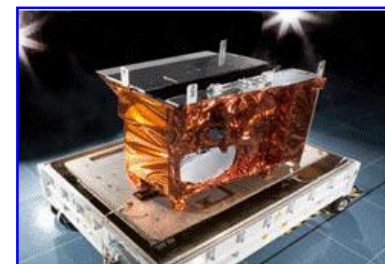
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Outline

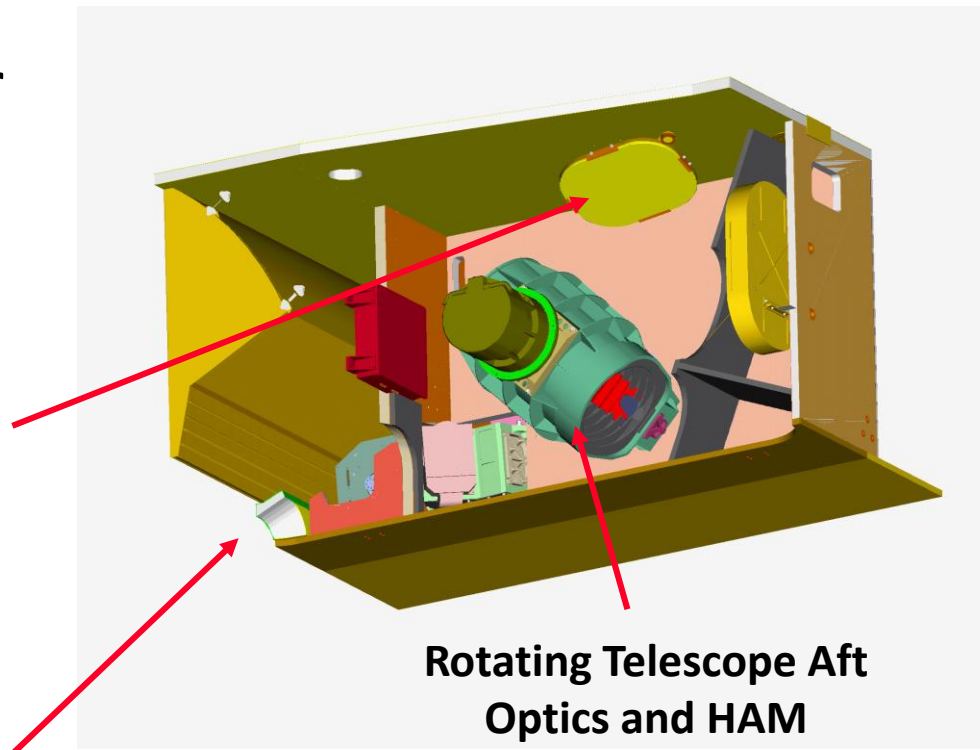
- **VIIRS Solar and Lunar Calibration**
 - Strategies and Activities
 - Methodologies
- **Performance and Discussion**
 - Changes and Updates (Improvements)
 - Comparison of Solar and Lunar Calibration
- **Future Efforts**

Solar and Lunar Calibration Strategies and Activities

**Solar Diffuser
Stability Monitor**



Solar Diffuser



**Rotating Telescope Aft
Optics and HAM**

SD:
each orbit

SDSM:
1 per orbit to 3 per
week

Moon:
8-9/year;
through SV port;
sector rotation;
fixed phase angle;
fixed high gain

Extended SV Port



- **15 reflective solar bands (RSB) including a DNB: 0.4-2.3 μm**
- 7 thermal emissive bands (TEB): 3.7-12.1 μm
- Moderate/Imaging bands: 750/375 m nadir spatial resolution
- Dual gain bands: M1-M5, M7, and M13

Solar and Lunar Calibration Methodologies

VIIRS Radiance (L) Retrieval: $L = F \cdot L_{PL} = F \cdot (c_0 + c_1 \cdot dn + c_2 \cdot dn^2) / RVS$

F : Calibration scaling factor derived from on-orbit calibration

c_i : Pre-launch calibration coefficients (quadratic algorithm)

RVS : Sensor response versus scan-angle

VIIRS Solar Calibration: $F_{SD} = \frac{L_{SUN}}{L_{SD,PL}}$

$$L_{SUN} \propto E_{SUN} \cdot BRDF(t) \cdot \tau_{SDS} \cdot \cos(\theta_{inc})$$

$$H(t) = \frac{dc_{SD} \cdot \tau_{SDSM}}{dc_{SUN} \cdot BRDF(t_0) \cdot \tau_{SDS} \cdot \cos \theta_{inc}}$$

L_{SUN} : Expected solar radiance reflected from SD panel

$L_{SD,PL}$: Retrieved solar radiance using pre-launch calibration coefficients

VIIRS Lunar Calibration: $F_{MOON} = \frac{I_{ROLO}}{I_{MOON,PL}} = \frac{I_{ROLO}}{\sum_{det,sam,scan} L_{MOON,PL} \cdot \Omega_B \cdot g / N_{SCAN}}$

I_{ROLO} : Lunar irradiance (integrated) provided by ROLO model

$I_{MOON,PL}$: Lunar irradiance retrieved using pre-launch calibration coefficients

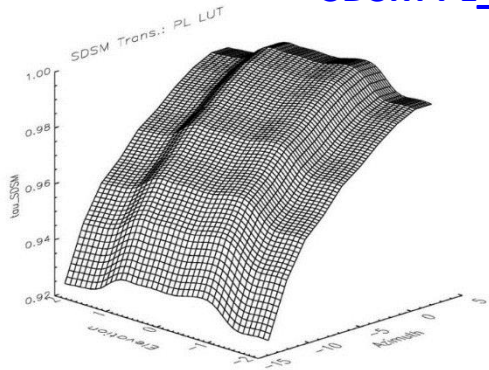
N_{SCAN} , Ω_B , g : number of scans, pixel solid angle, aggregation factor

Changes and Updates (Improvements)

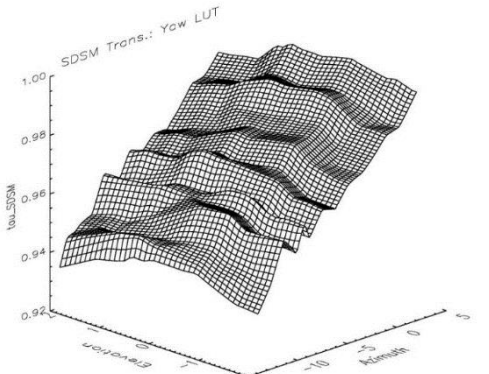
- **SD and SDSM Screen Transmission**
 - Pre-launch characterization
 - On-orbit yaw maneuvers
 - Yaw + regular on-orbit data
- **Correction for Solar Vector Error**
 - Consistently reprocessed SDR for NASA science research community
 - Different impact for VIS/NIR and SWIR
- **Modulated RSR (relative spectral response)**
 - Due to strong wavelength-dependent optics degradation
 - Different impact for solar and lunar calibration, and EV data
 - Large effect for DNB (broad bandwidth: 500-900 nm) calibration

SD and SDSM Screen Transmission (look-up table)

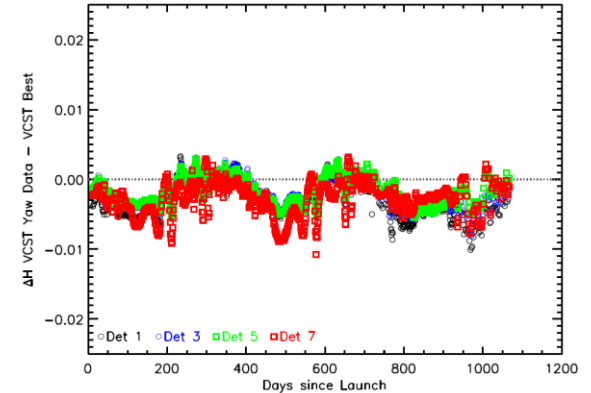
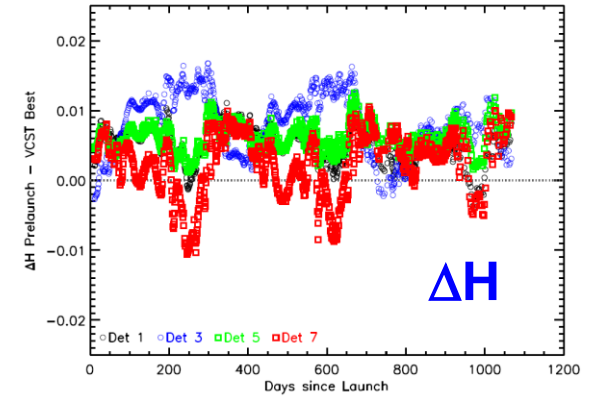
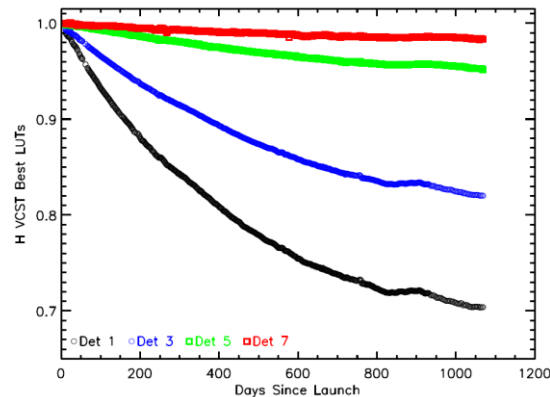
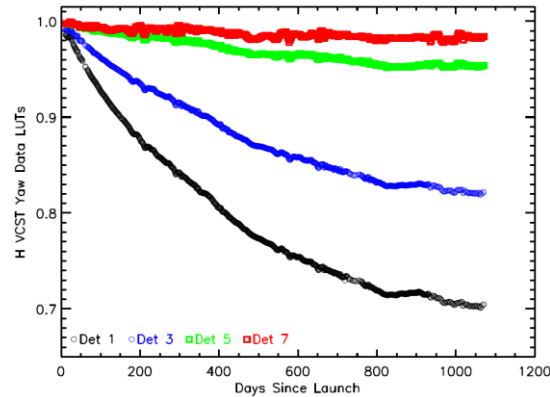
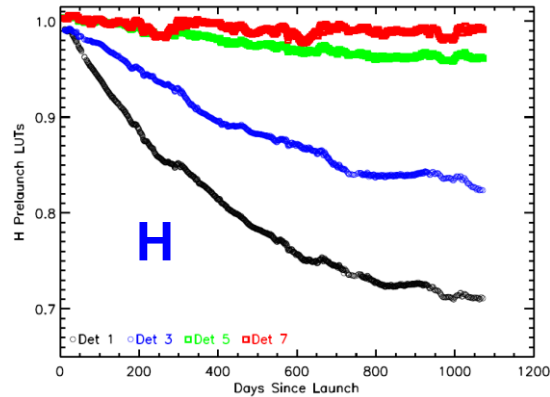
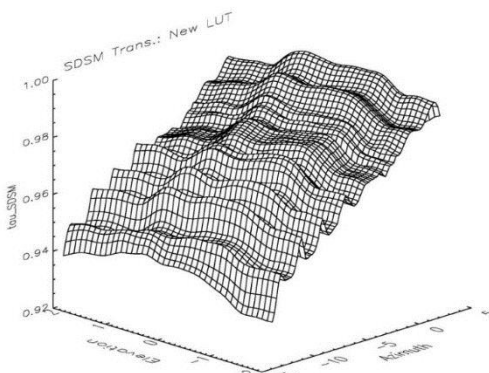
SDSM PL_LUT



SDSM Yaw_LUT



SDSM New_LUT



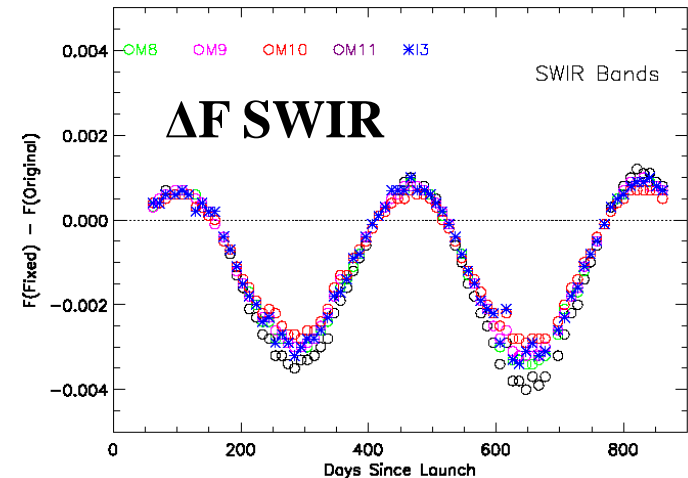
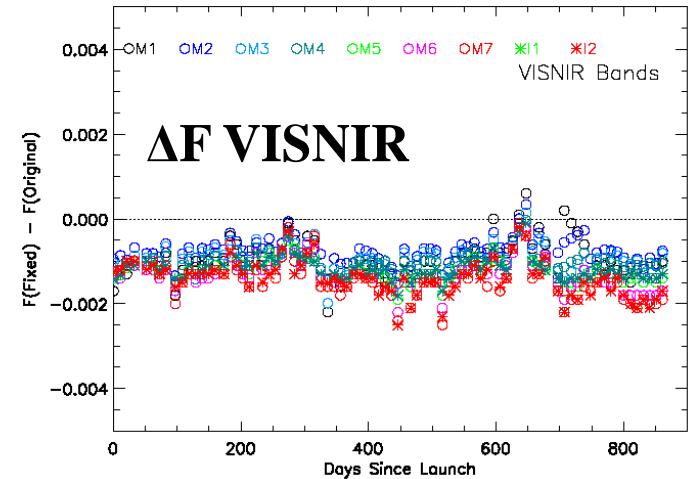
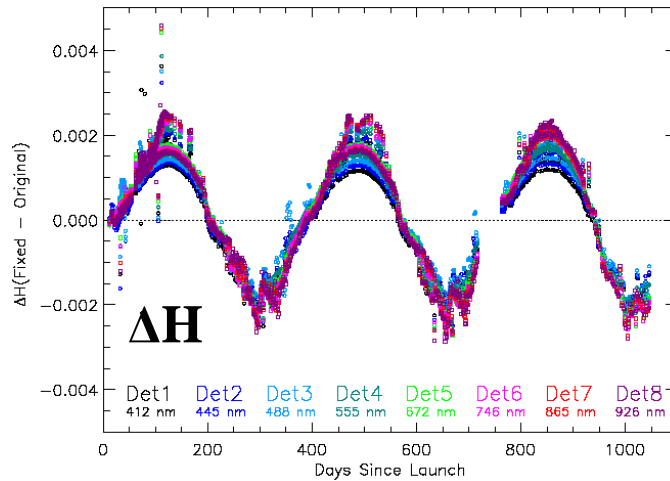
Improved H => better quality of F

Similar improvements to SD screen transmission

0.41, 0.48, 0.67, 0.86 μm

Correction for Solar Vector Error (in SDR Geo Library)

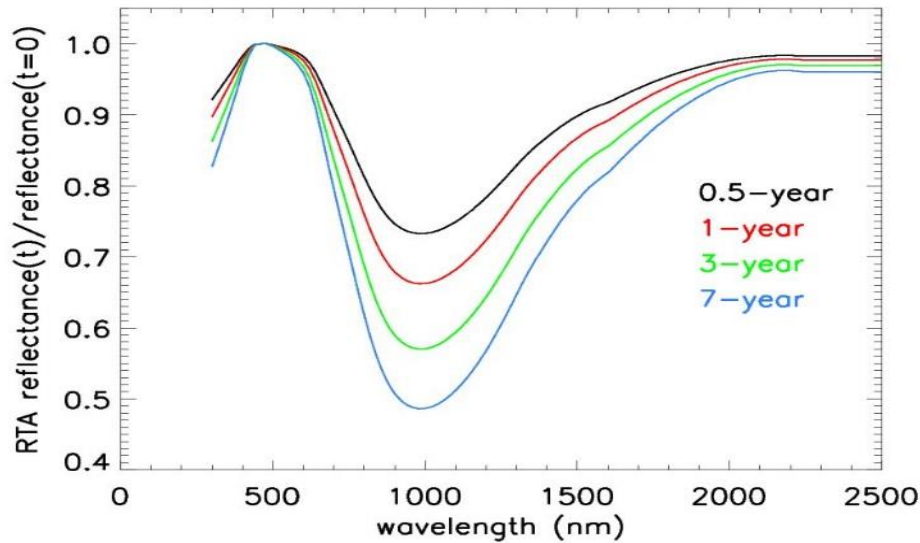
- A mismatch of ECI (Earth-Centered Inertial) frames when computing the transformation to spacecraft frame library leads to $\sim 0.2^\circ$ error in the solar angles used in the RSB radiometric calibration.
- Different impact for VIS/NIR and SWIR bands



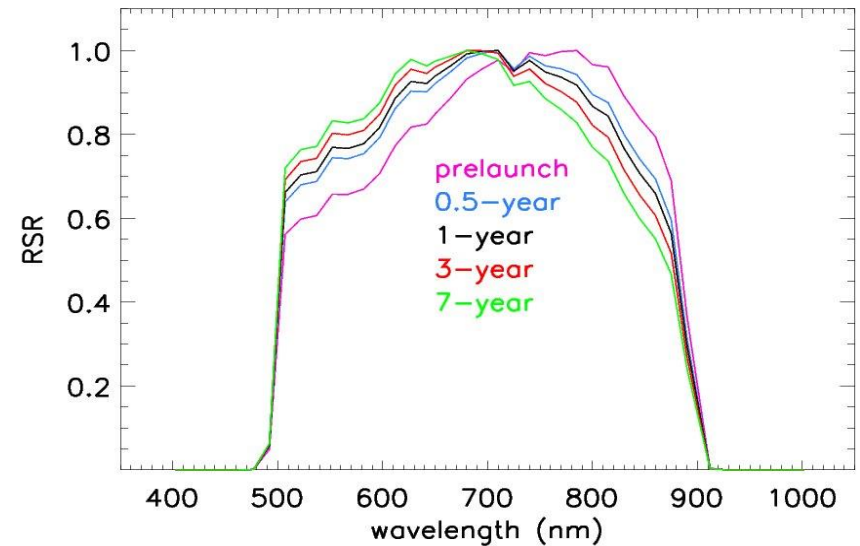
- The $\cos \theta_{SD}$ factor is used in both H- and F-factor calculations.

Development and Update of On-orbit Modulated RSR

λ dependent optics degradation



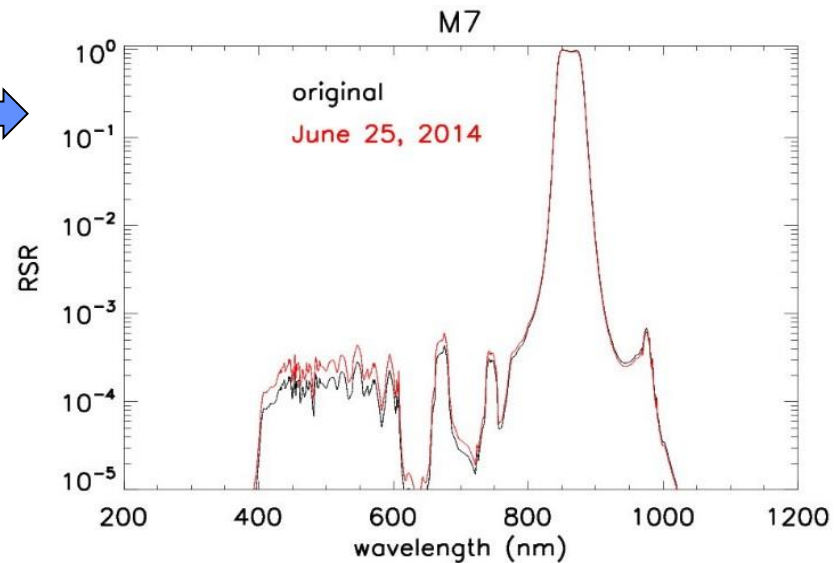
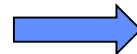
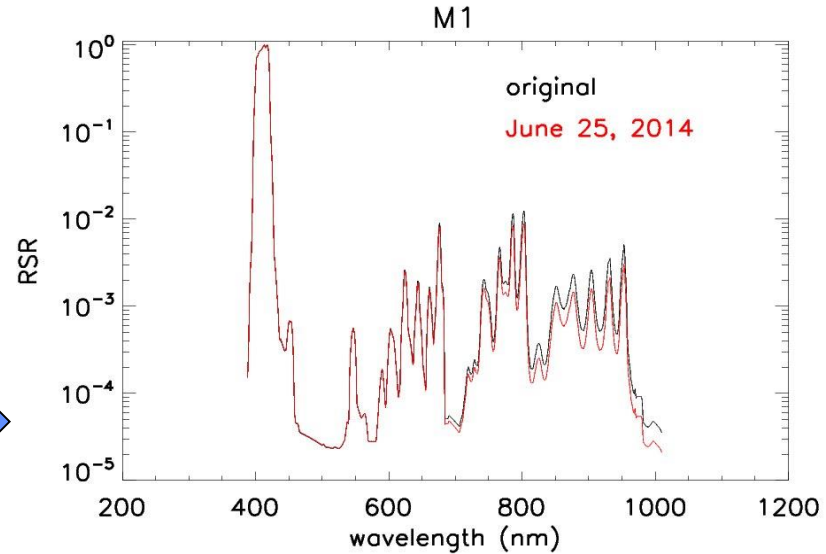
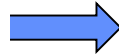
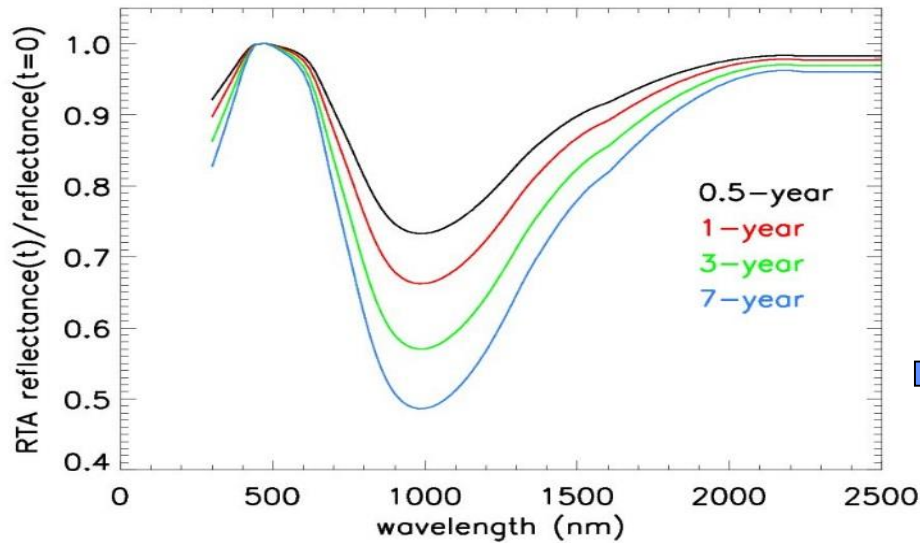
Modulated RSR



large impact on DNB (broad bandwidth)

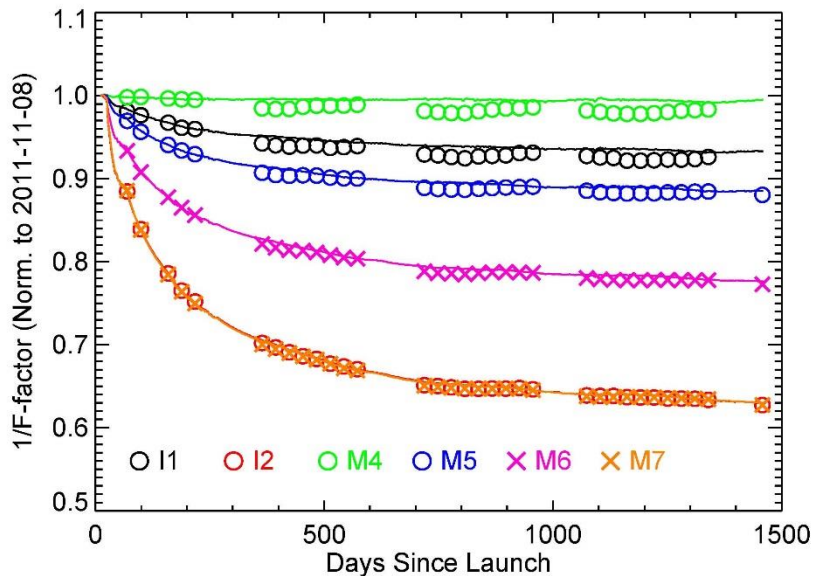
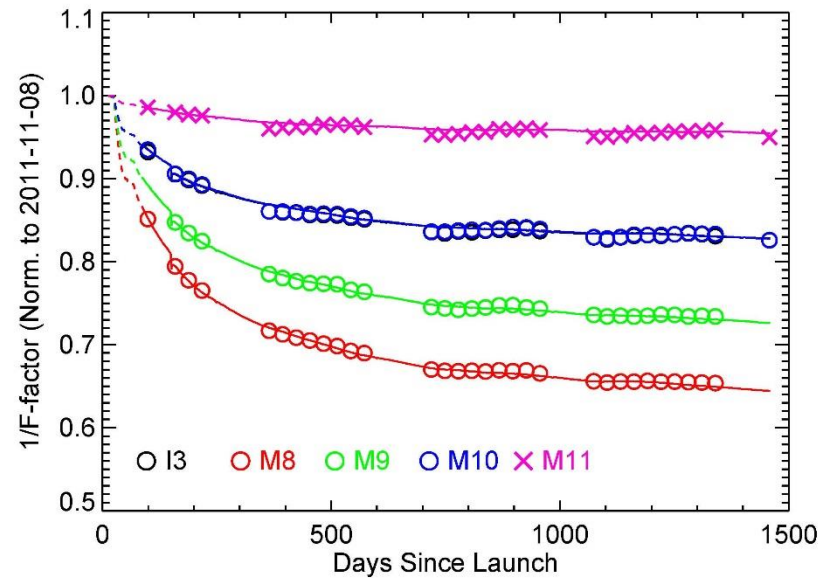
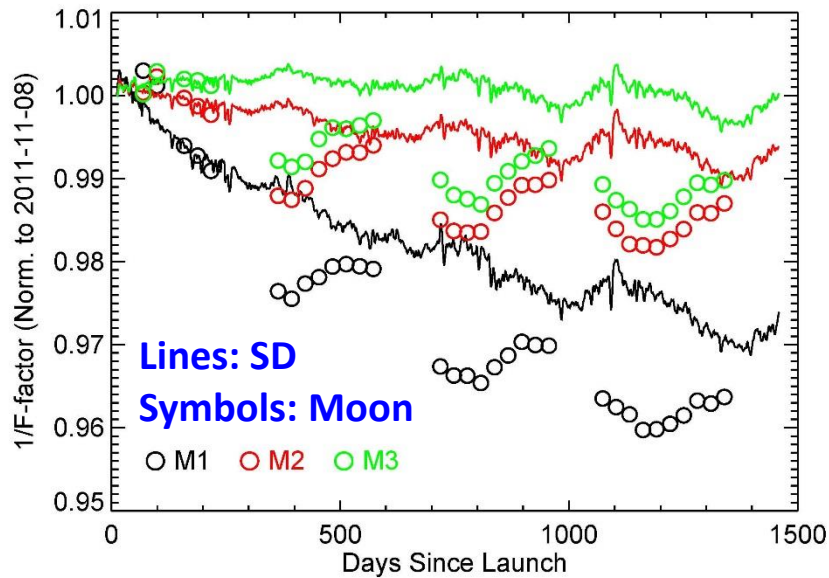
Development and Update of On-orbit Modulated RSR

λ dependent optics degradation



Small impact on bands with narrow bandwidths and small OOB responses

Comparison of Solar and Lunar Calibration



SD and lunar calibration made at the **“same”** angle of incidence (AOI)

Large changes in NIR/SWIR response due to telescope mirror degradation

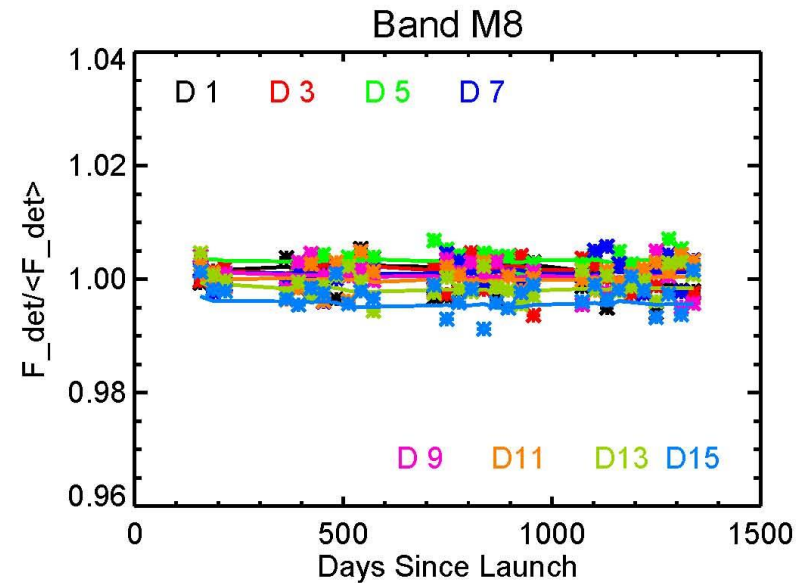
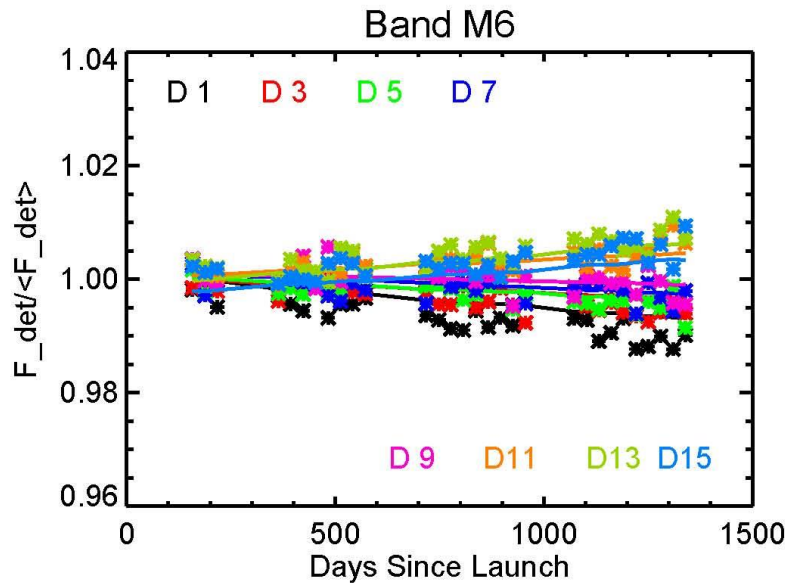
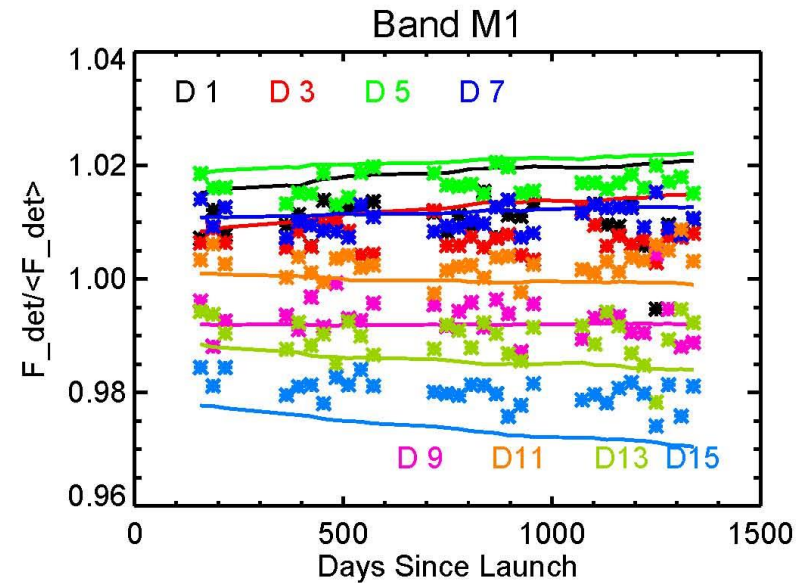
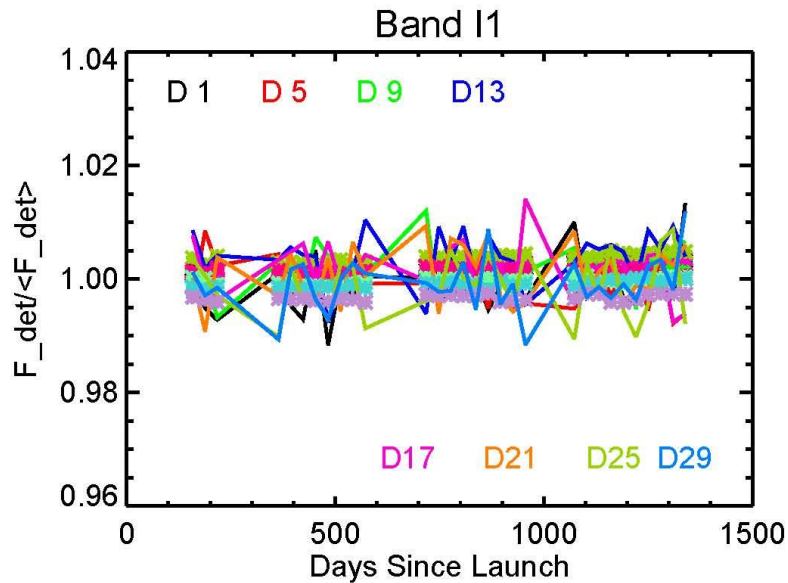
Modulated RSR applied

M1	0.41
M2	0.45
M3	0.49
M4	0.56
I1	0.64
M5	0.67
M6	0.75
I2	0.87
M7	0.87
M8	1.24
M9	1.38
I3	1.61
M10	1.61
M11	2.25

Future Efforts

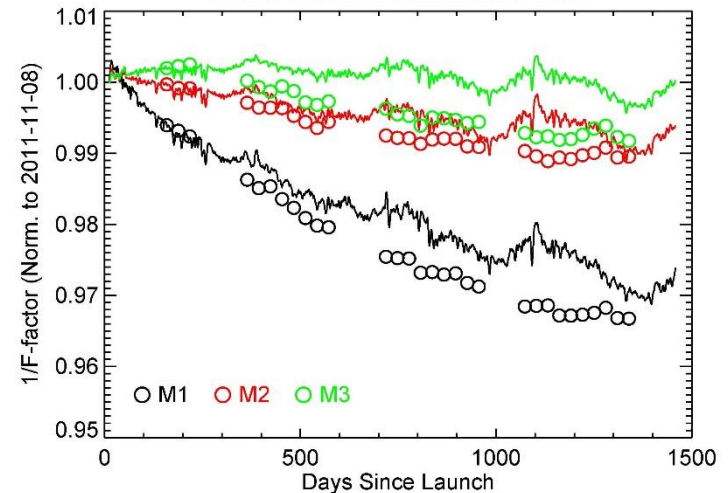
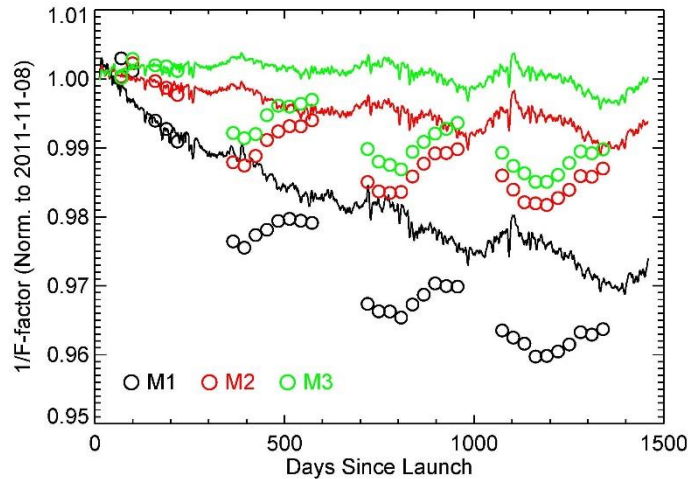
- **Combine SD and Lunar Calibration for Improved SDR LUTs**
 - SD and lunar observations are made at the same AOI
 - Remove potential impact due to SD degradation (SDSM and SD degradation uniformity)
- **Use Lunar Observations to and Characterize and Reduce Detector to Detector Calibration Differences**
 - MODIS experience and lessons applied
 - Small differences in a few VIIRS spectral bands
- **Improve Lunar Calibration**
 - Absolute (effort by NIST/USGS and by GSICS)
 - Relative (response trending and calibration inter-comparison)

Detector to Detector Calibration Differences

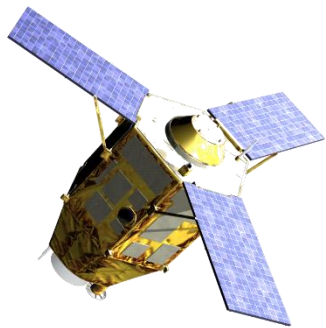


Approaches for Lunar Calibration Improvements

With an empirical libration correction



Impact due to lunar phase angles



Pleiades: POLO

